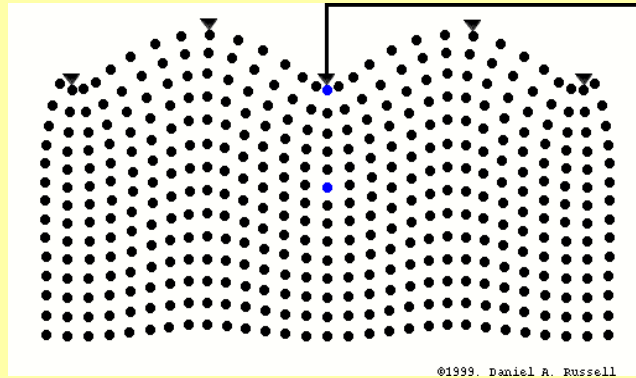


Ground-roll Inversion for S-wave Velocity: Meteor Crater, Arizona



by
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Supervisor: Dr. Robert R. Stewart

Allied Geophysical Laboratories (AGL) Meeting and Research Day

University of Houston
29th April, 2011



Overview

Motivation and Objective

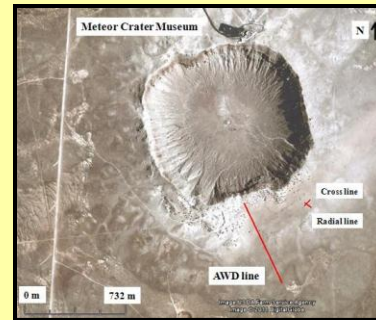
- Near-surface problems

Methodology

- Frequency based ground-roll inversion method

Barringer (Meteor) Crater, Arizona

- Geological Setting
- Seismic experiments



Results and Interpretations

- S-wave velocities (V_S) and identifying geological features

Applications and Future works

- Density prediction
- S-wave statics
- 3D Tomography

Motivations and Objectives

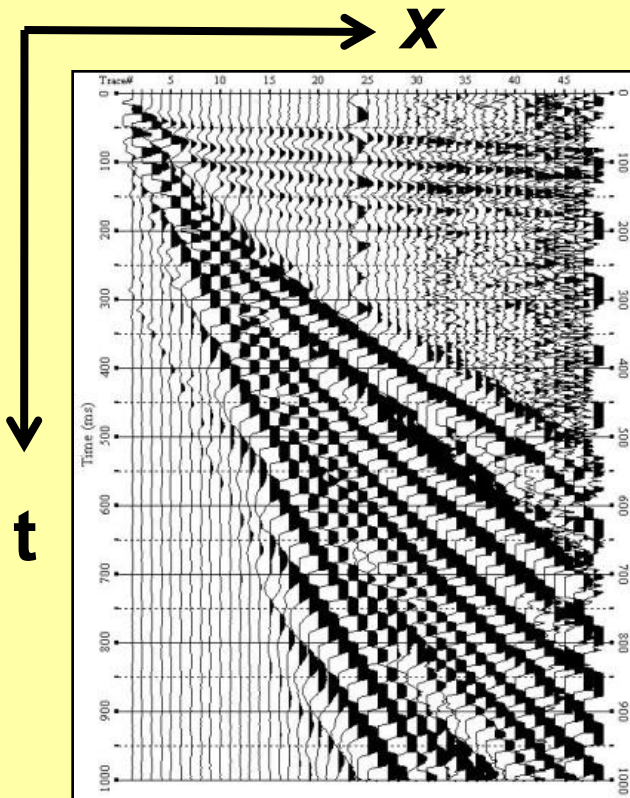
- **Study of the near-surface velocity structure**
 - **Low-velocity, unconsolidated layers**
 - **Near-surface effects propagate deeper**
- **A good estimation of V_S required for multicomponent seismic analysis**
- **Identify geological features using estimated V_S variations**
 - **Thickness of a low-velocity, unconsolidated layer at Meteor Crater, AZ**
 - **Correlation of the results with known information**
- **Time delays in seismic waves (Statics problem)**

$$t_{\text{Statics}} = \sum \Delta z_i / V_i$$

Methodology

- Multichannel Analysis of Surface Waves (MASW)
 - Generation of dispersion curves (phase velocity versus frequency plots)
 - Inversion of dispersion curves to estimate V_s

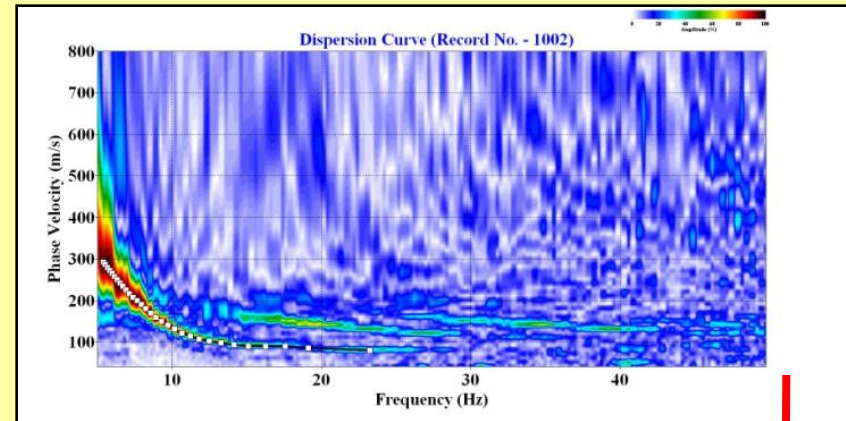
(Park et al., 1998; ibid 1999; Xia et al., 1999)



Raw shot gather: La Marque, TX

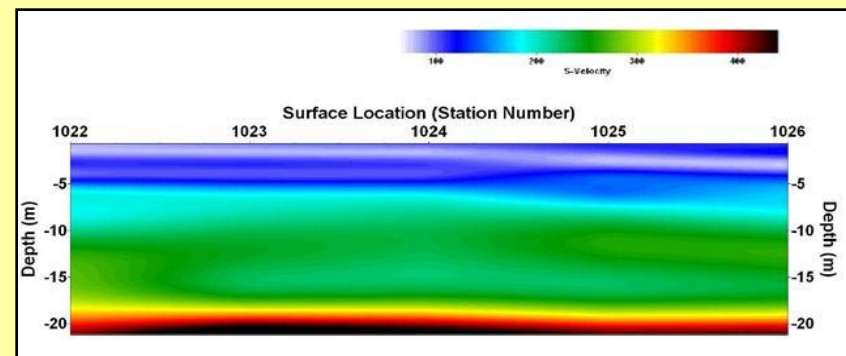


Phase velocity (m/s)



Frequency (Hz)

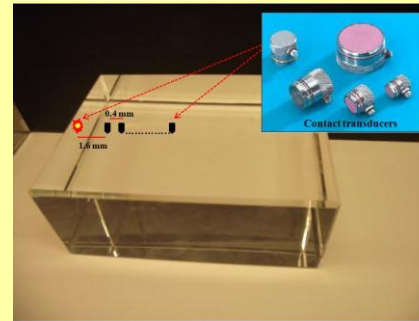
Depth (m)



Different Seismic Data Sets

Modeling data

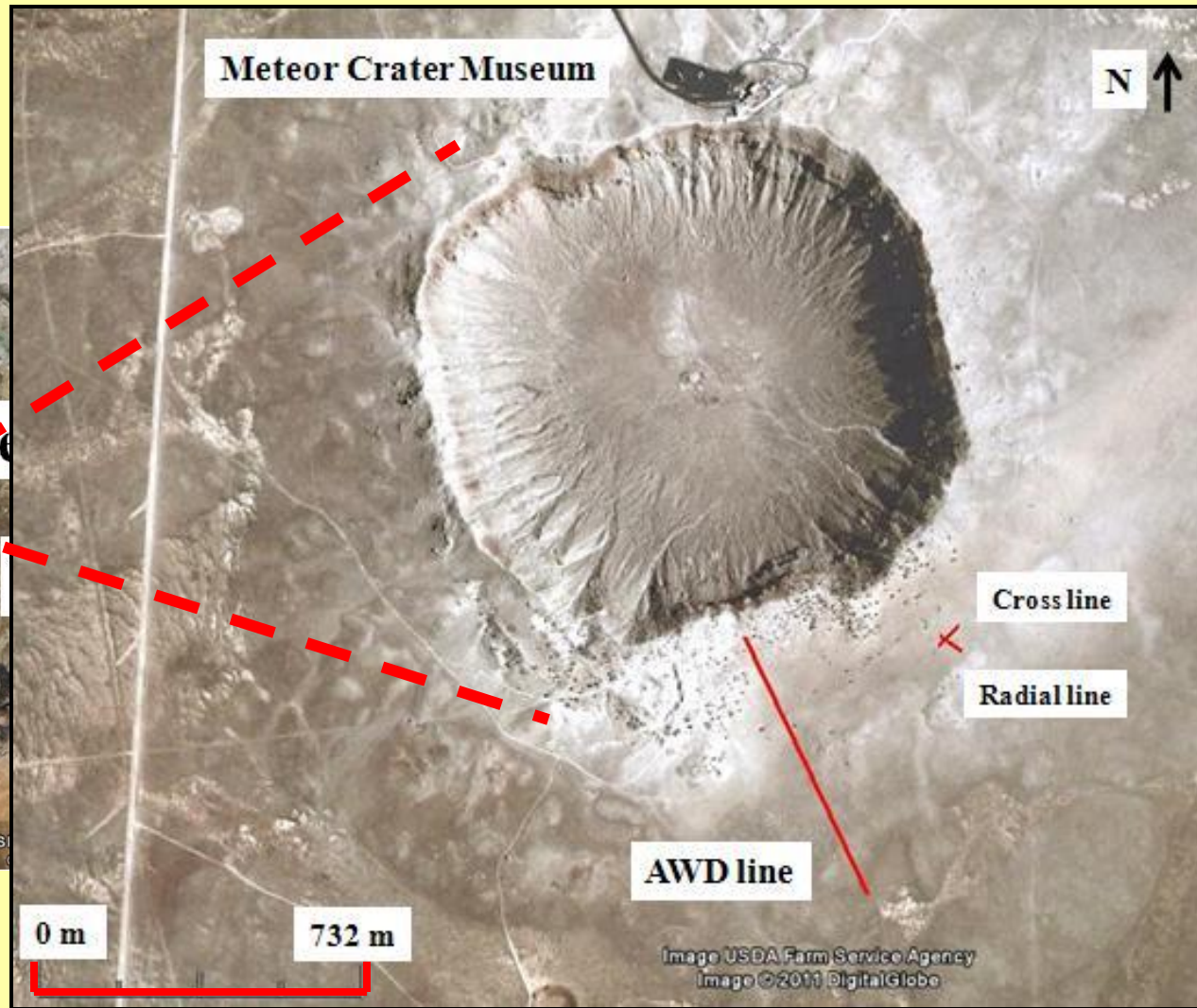
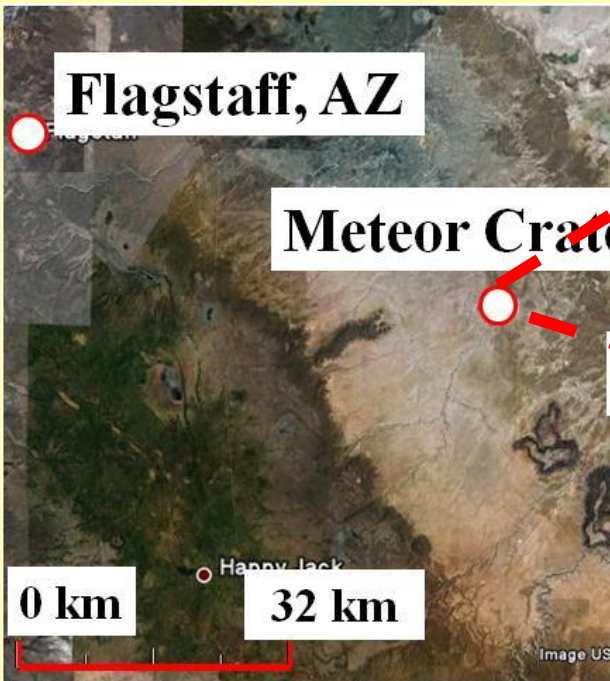
- Numerical modeling:
Finite-difference elastic modeling (Manning, 2007)
- Physical modeling:
Ultrasonic measurements on
glass model



Field data

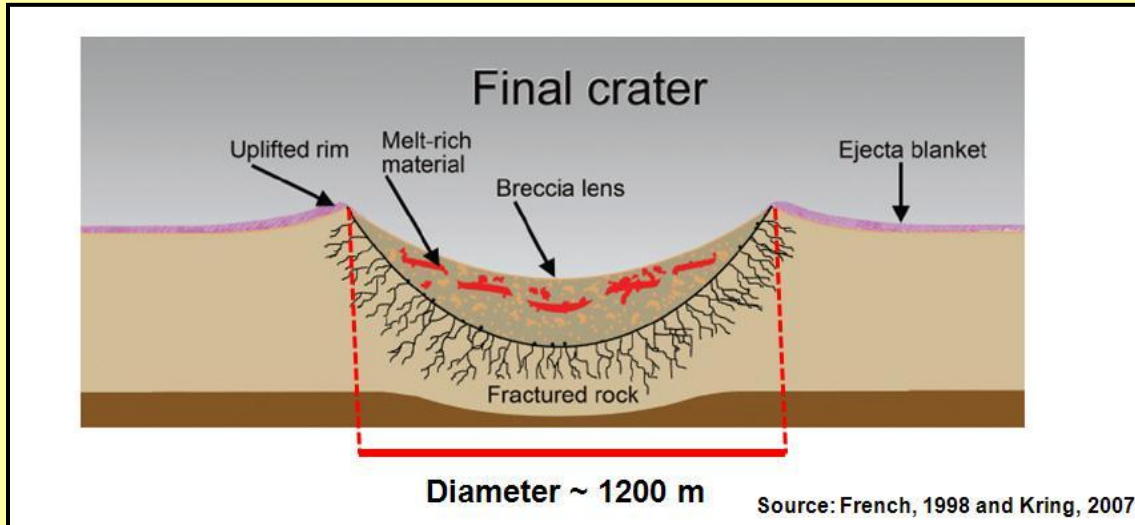
- **Barringer (Meteor) Crater, AZ**
- La Marque Geophysical Observatory, TX
- Yellowstone-Bighorn Research Association (YBRA), MO

Barringer (Meteor) Crater, AZ



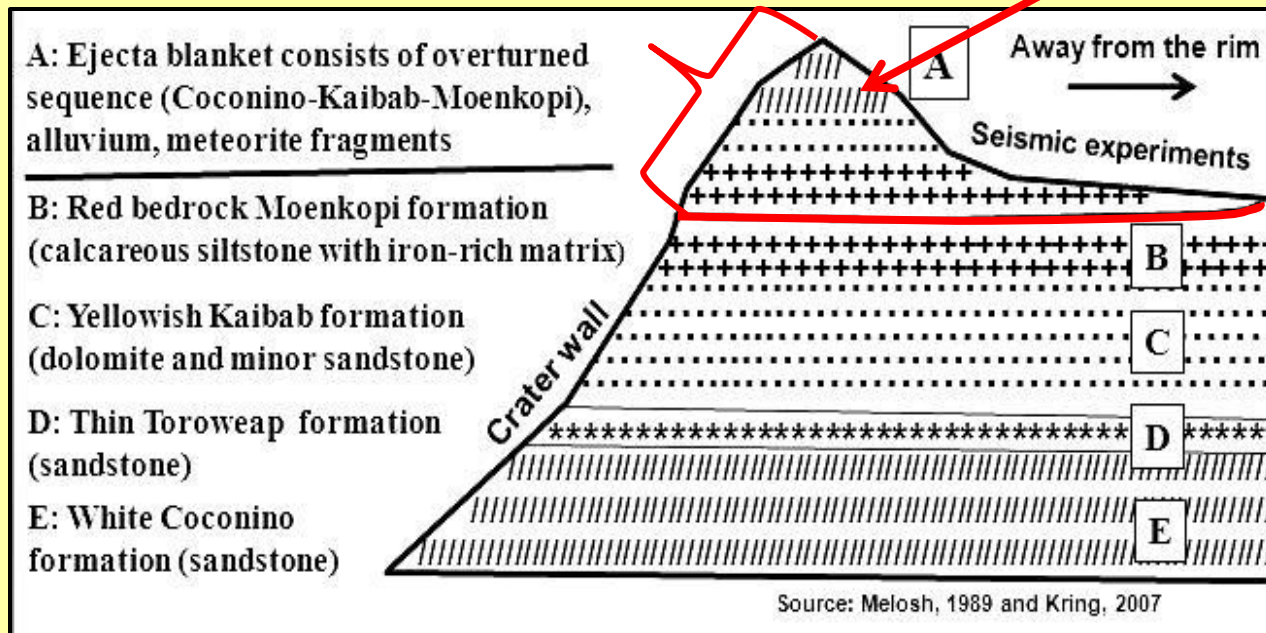
Source: Satellite image from Google Earth

Meteor Crater: An Unique Astrobleme



- Meteorite Craters in Texas:**
- Odessa Crater
 - Sierra Madera Crater

Overtured sequence



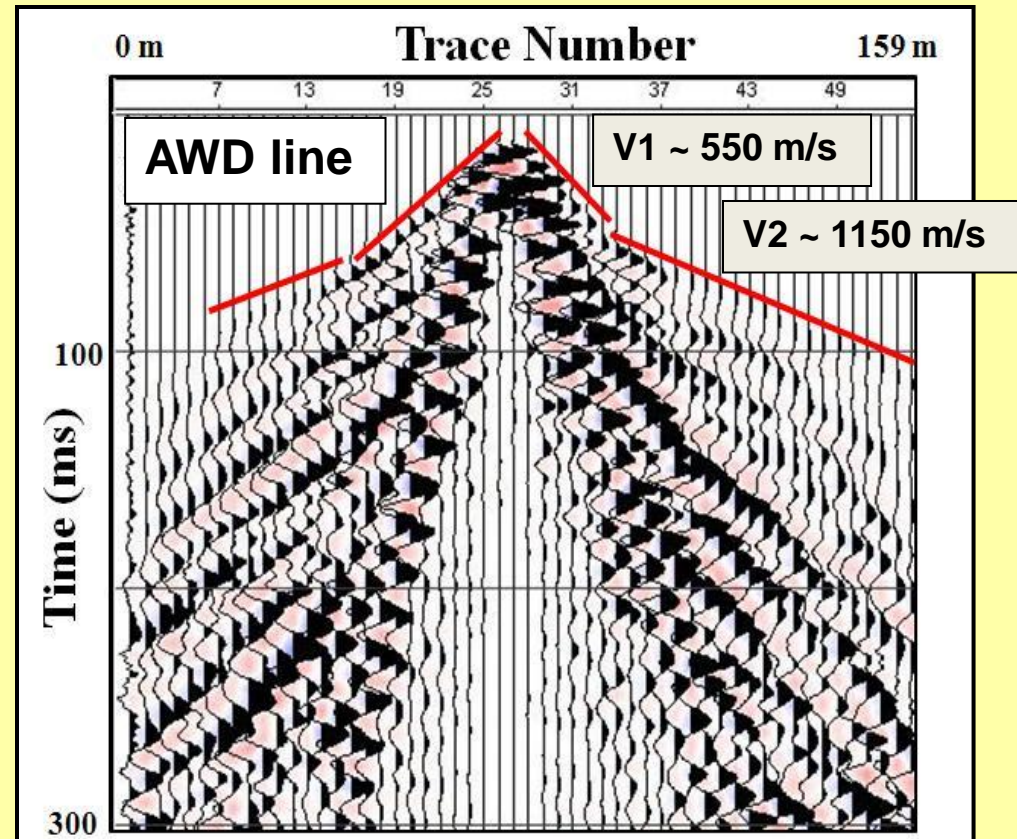
Seismic Experiments at Meteor Crater

- Source interval: 2-3 m
- Receiver interval: 2-3 m
- Source type: 10 lb sledgehammer, Accelerated Weight Drop

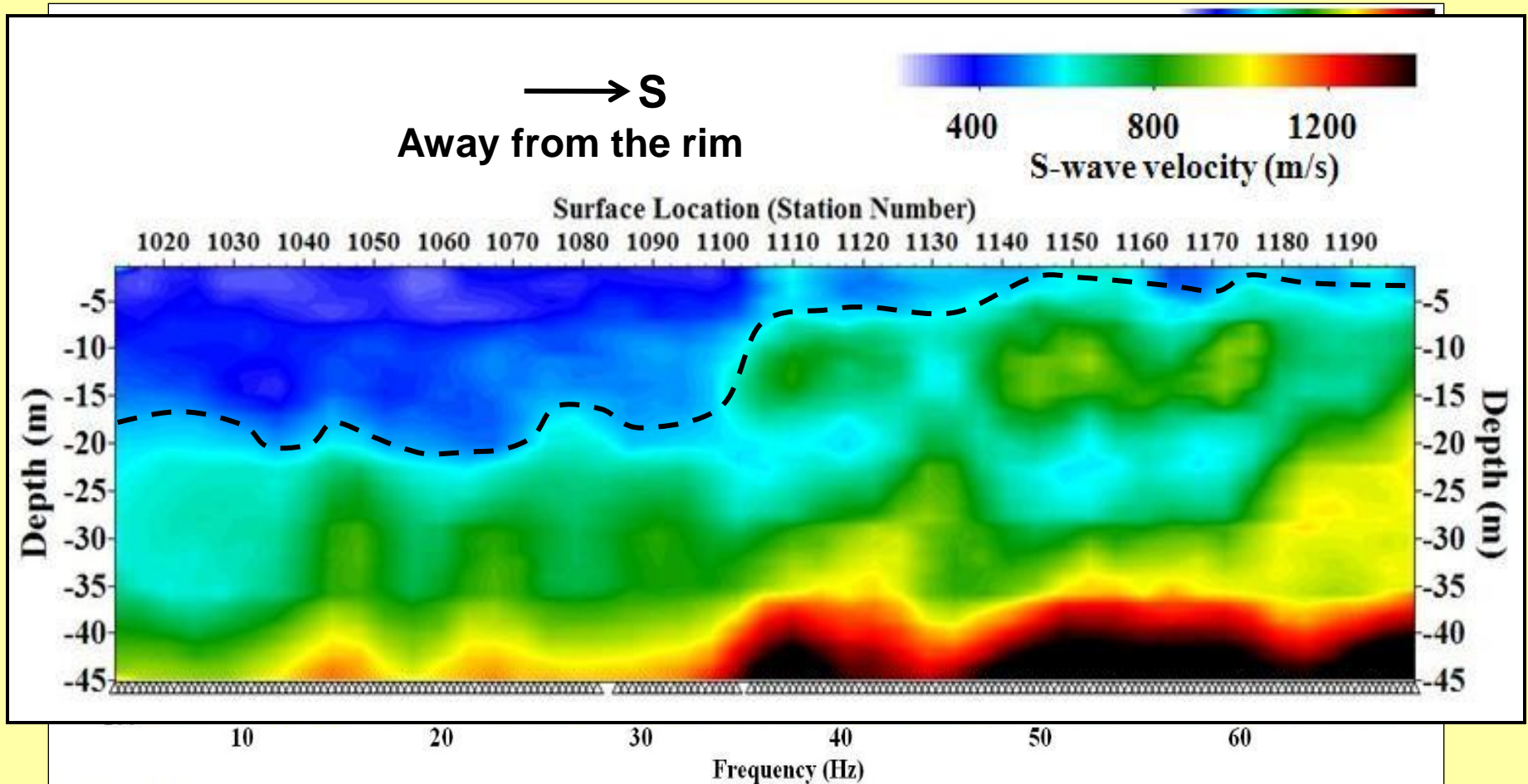
Ultrasonic Measurements



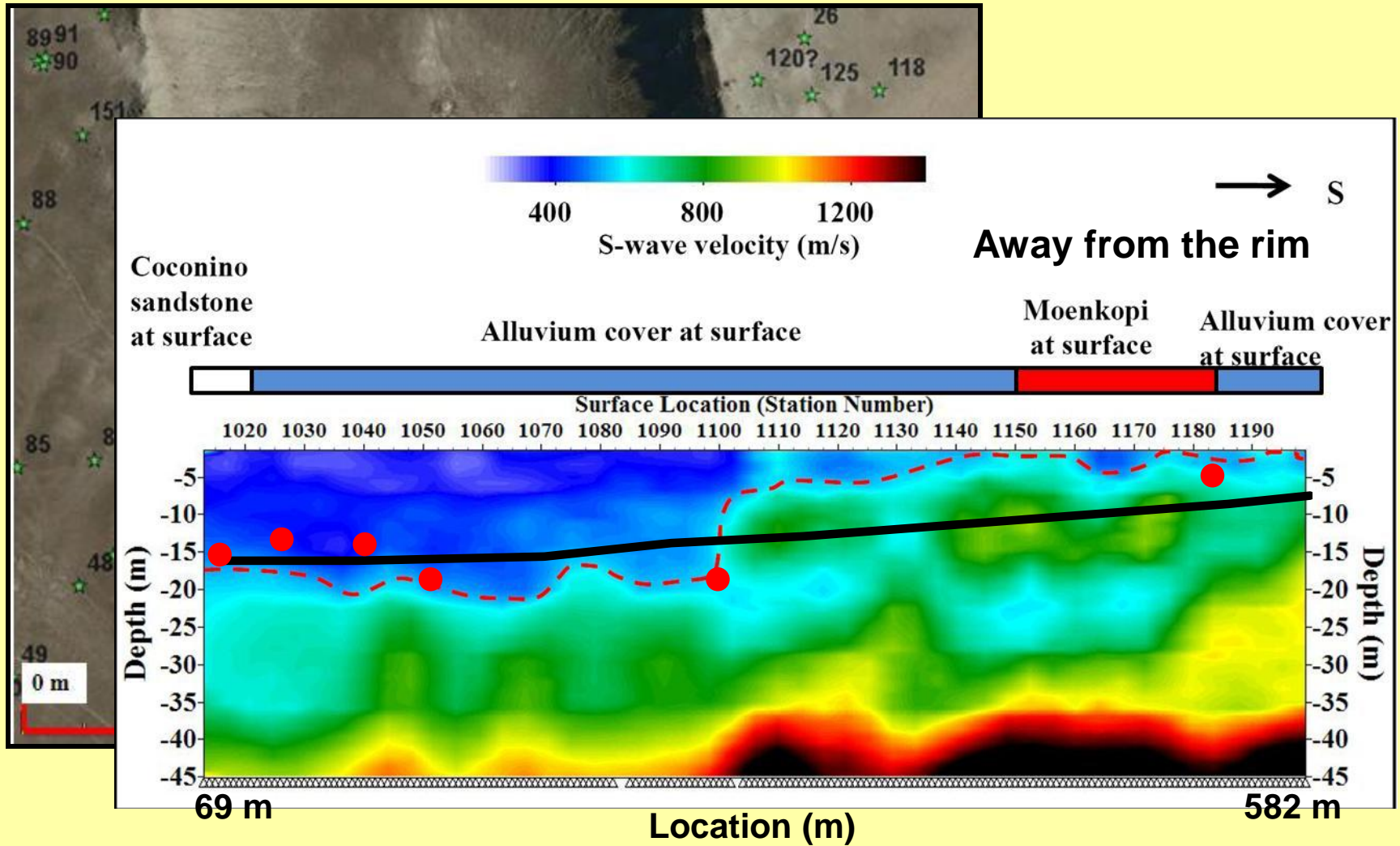
Rock Formation	Range of P-wave velocity (m/s)
Moenkopi 1	815 ± 33
Moenkopi 2	1255 ± 106
Moenkopi 3	1570 ± 89
Kaibab 1	3705 ± 81
Kaibab 2	2560 ± 210



Result: Dispersion curve and S-wave velocity



Interpretation: Meteor Crater



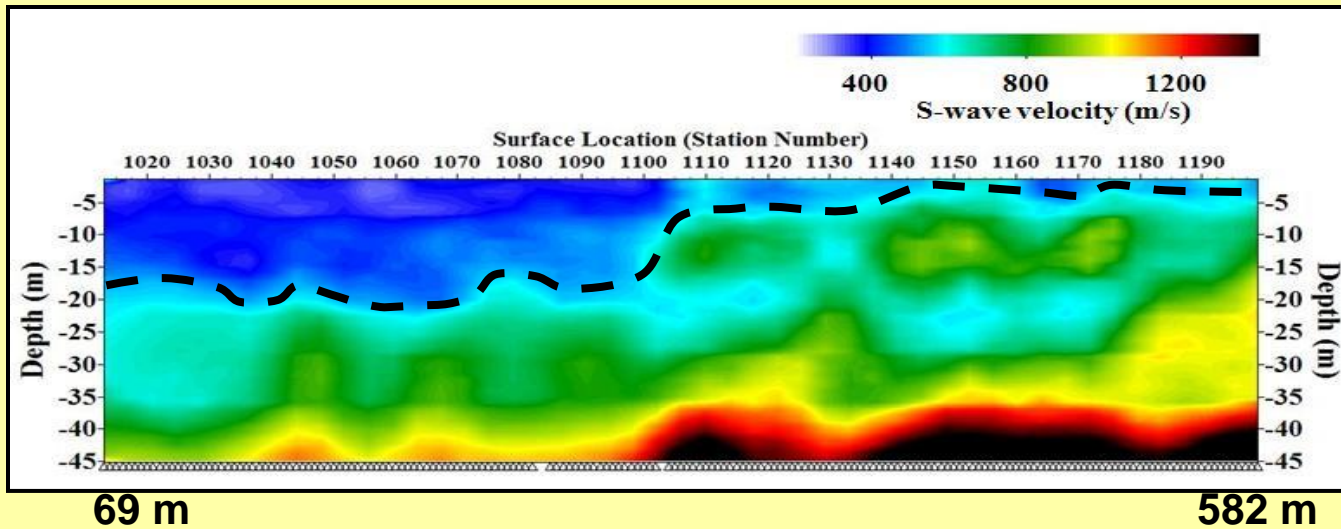
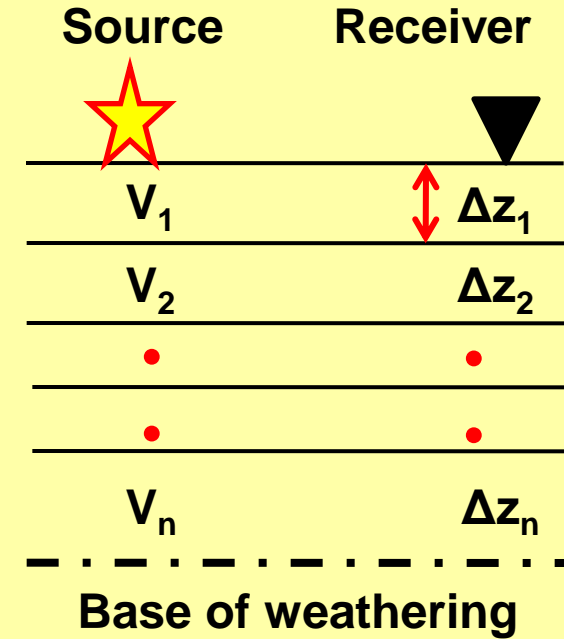
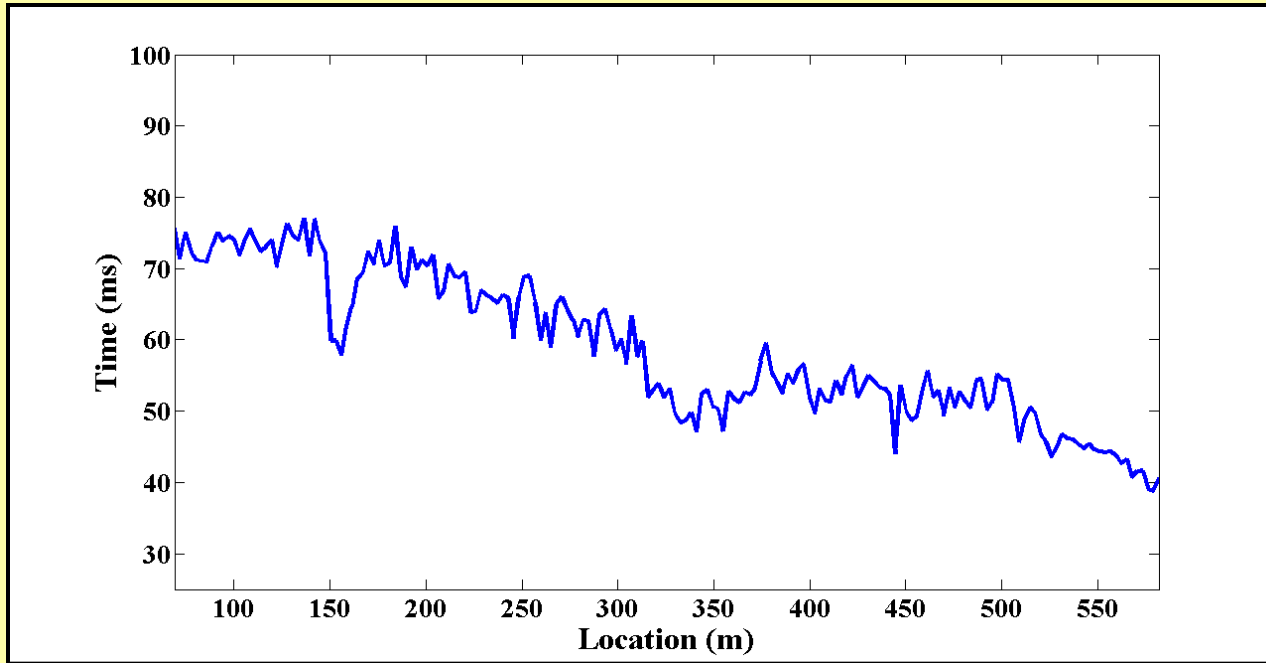
Conclusions

- Ground-roll inversion is a simple, robust, noninvasive, and low-cost method
- 2D S-wave velocity structures can be obtained
- A broad range of S-wave velocities are estimated:

Vs (m/s)	Numerical Model	Physical model	YBRA, MO	La Marque, TX	Meteor Crater, AZ
Known	500-740	3400	300-900	N/A	N/A
Ground-roll inversion	450-800	3200-3500	300-1000	100-500	200-1000

- Geological structures are identified based on the estimated S-wave velocity variations

Application: S-wave Low-velocity Statics



$$t_{\text{Statics}} = \sum \Delta z_i / V_i$$

69 m

582 m

Application: Density from S-wave velocity

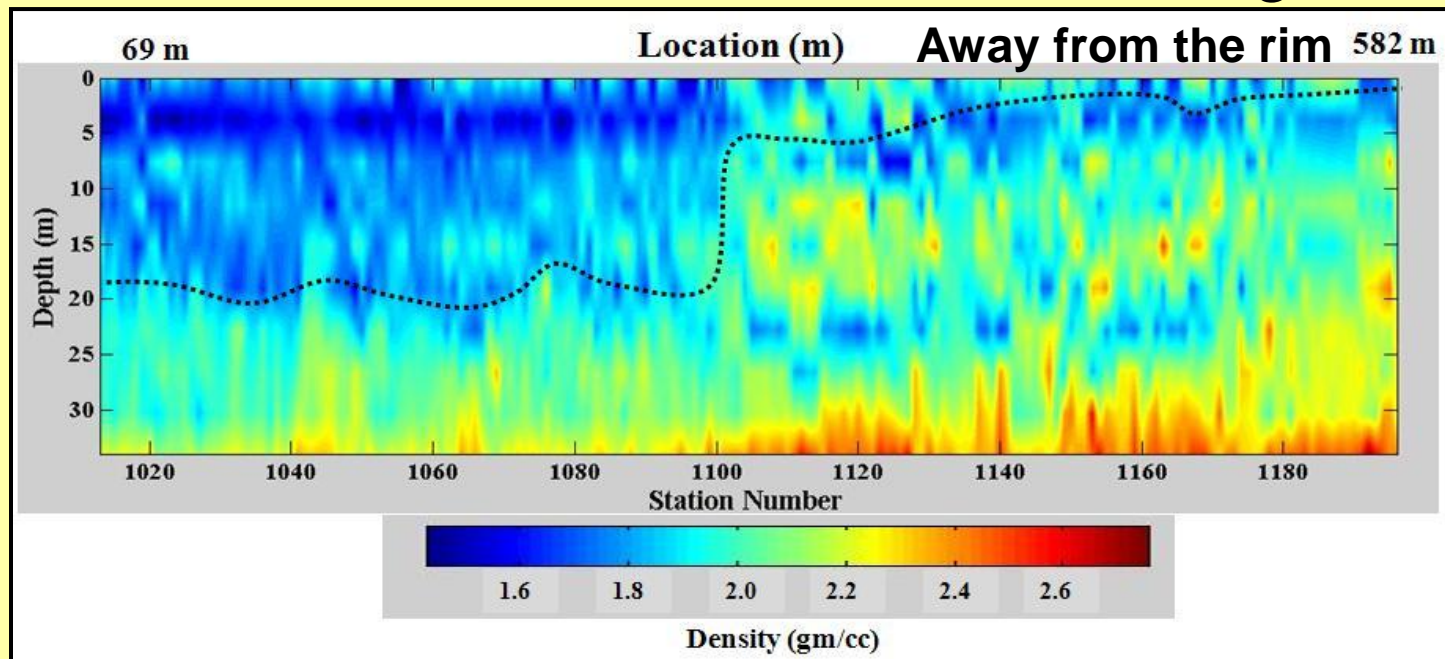
Gardner's Equation for S-wave: $\rho = 0.37 V_S^{0.22}$

Where, ρ in gm/cc and V_S in ft/s - (Potter and Stewart, 1998)

Target units	Bulk density (gm/cc) from drill cores
Ejecta blanket	1.87-2.17
Moenkopi	2.19-2.48

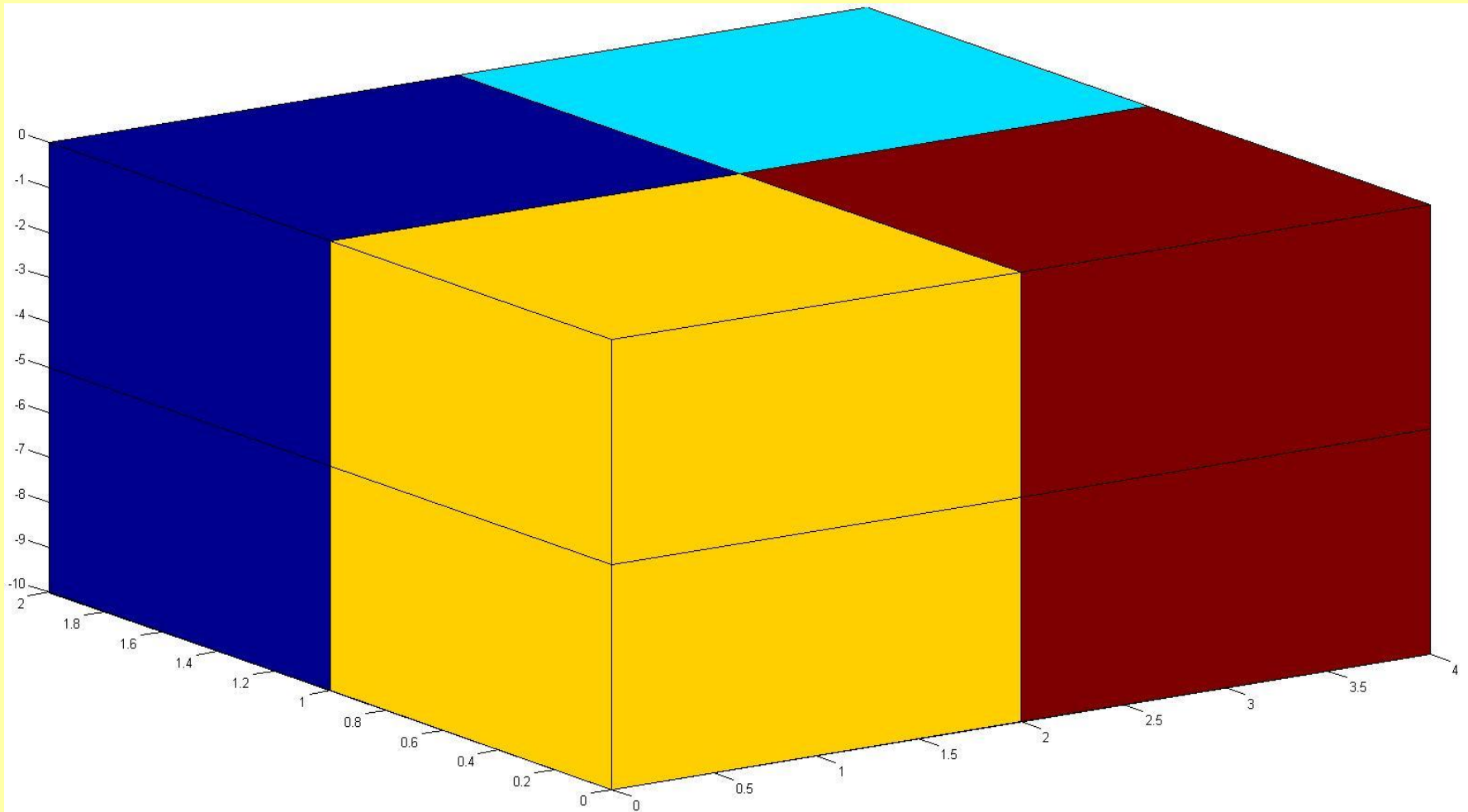
(Walters, 1966, Watkins *et al.*, 1966, and Kring, 2007)

→ S



Future work: 3DTomography

- Construction of a picture of the internal values of an object from sum of those values
- Finding an integrand from various integral values (Stewart, 1991)



Acknowledgement

- **Allied Geophysical Laboratories (AGL)**
- **My colleagues at AGL**
- **Dr. D. A. Kring (Lunar and Planetary Institute)**
- **Meteor Crater Museum**
- **Kansas Geological Survey**



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