

ABSTRACT

We investigated the shallow subsurface at Barringer (Meteor) Crater, Arizona using seismic and gravity techniques. We found compressional (P)-wave velocities of 450-2500 m/s for a 55 m deep section from seismic refraction analysis. The low uppermost P-wave velocity layers thin Trace Number away from the crater rim (toward the south). Shear (S)-wave velocities (estimated from ground-roll inversion) vary from 200-700 m/s for the top 16-20 m increasing to 900-1000 m/s at 38 m depth. The prominent change in S-wave velocities (at around 500-600 m/s) is interpreted as the transition from the low-velocity ejecta blanket (a sheet of debris thrown out of the crater during the impact) to the bed-rock Moenkopi sandstone. This S-wave transition takes place at a depth range of 12-20 m near the crater rim with a thinning away from the rim. This consistent Pand S-wave low velocity structure is interpreted as the ejecta blanket. Near-surface reflection seismic analysis provided relatively deeper information indicating at least four sub-horizontal reflectors intersected by a number of faults. Staking velocity analysis estimates P-wave velocities up to 3500 m/s (for the top 500 ms) showing a similar P-wave thinning structure (especially for top 200 ms). Ultrasonic measurements on hand samples provide a range of P-wave velocities of 800-1600 m/s for the Moenkopi and are consistent with seismic refraction results.

Objectives

- To understand seismic wave propagation through brecciated materials
- To estimate thickness of ejecta blanket (a sheet of debris thrown out of crater during meteorite impact)
- To characterize the near-surface physical properties
- To develop general survey methodologies to image a highly complex near-surface
- To image near-surface reflectors and faults

Barringer (Meteor) Crater, Arizona



Figure 1. Schematic diagrams showing the creation of a simple crater (1-3), and the present day stratigraphy (4) at Meteor Crater (a simple crater).

Seismic Survey

Table 1: Seismic acquisition parameters

Seismic Line	AWD
Source Type	40 kg Accelerated weight drop (AWD)
Receiver Type	Planted vertical geophone
Source Interval (m)	3
Receiver Interval (m)	3
Total Receivers	216



Figure 2. Image of LiDAR data for southern portion of Meteor Crater. Red line represents the seismic line used in this paper.

32: Debris from Kaibab formation A2: Debris from Moenkopi formation A1: Moenkopi formation **B1: Kaibab formation**

Coconino and Toroweap formations

Ultrasonic Measuements

 Table 2: Ultrasonic measurements of
Moenkopi hand specimens showing P-wave velocities.

Rock formation	P-wave velocity (m/s)
Moenkopi 1	815 ± 33
Moenkopi 2	1255 ± 106
Moenkopi 3	1570 ± 89



(Shoemaker *et al.*, 1974 and Kring, 200



near-surface S-wave velocity structure.

dispersive ground-rolls (or Rayleigh waves).





AWD line and locations of the drill-holes (Roddy et al., 1975). actual elevation).





Figure 9. Transition depths (from the ejeta Figure 8. Overlay image of the blanket to bed-rock Moenkopi) obtained from different analyses and drill-holes (ploted from



• The tickness of the ejecta is also estimated.

• P-wave NMO velocity structure shows a similar thinning pattern.

• Few sub-horizontal reflectors intersected by number of faults are identified using near-surface seismic reflection analysis.

 Near-surface geophysical methods as outlined above may be useful in studying other sites with similar complex brecciated structures.

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